

## SIO 231 Geoelectromagnetism

Geomag Homework # 3,

Due February 13, 2024

**Q1** On the class web set you will find a file called chaos\_7\_9\_av\_lm containing Schmidt normalized Gauss coefficients for the CHAOS7-9 internal geomagnetic field averaged over the time interval 1997.0 to 2022.1.

Calculate and plot the geomagnetic spatial power spectrum  $R_l$  at Earth's surface (radius  $a = 6371.2\text{km}$ ), at two additional radii of your choice one above and one below  $r = a$ , and at the core mantle boundary ( $r = c = 3485\text{ km}$ ). Use the spectrum at Earth's surface and its rate of fall off to estimate the radius of Earth's outer core. Redo the calculation excluding the  $m = 0$ , zonal Gauss coefficients and comment on the origin of the differences.

**Q2** In class we discussed the possibility that paleomagnetic secular variation could be described by a Giant Gaussian Process where time variations are ascribed a statistical representation in terms of each Gauss coefficient for degree and order  $l = 1 - 14$ . One GGP published by Lisa Tauxe and Dennis Kent in 2004 has the feature that the variances for the  $g_l^m$  and  $h_l^m$  depend on whether  $l - m$  is even or odd which correspond to SH functions that are respectively symmetric or antisymmetric about the equator. Slide 27 gives

$$\sigma_l^{m2} = \frac{(c/a)^{2l} \alpha^2}{(l+1)(2l+1)} \quad \text{for } l - m \text{ even}$$

$$\sigma_l^{m2} = \frac{(c/a)^{2l} \alpha^2 \beta^2}{(l+1)(2l+1)} \quad \text{for } l - m \text{ odd.}$$

with  $\alpha = 7.5\mu\text{T}$  and  $\beta = 3.8$ . All the mean values  $\bar{g}_l^m$  and  $\bar{h}_l^m$  are zero except for the axial dipole  $\bar{g}_1^0 = -18\mu\text{T}$ . Calculate and plot the spatial power spectrum up to degree and order 14 for this GGP. How do things change if the parameter  $\beta$  is modified to have value  $\beta = 1$ ? And what would you expect if the  $\sigma_l^{m2}$  are allowed to vary within each SH degree  $l$ ?

**Q3** Also on the class web set you will find a file containing MAGSAT observations of  $B_\theta$  and  $B_r$  as a function of colatitude and radial distance from Earth's center. Plot the measurements as a function of  $\theta$ . Use these to make a least squares estimate of  $i_1$  and  $e_1$  for this pass as described in Lecture 9. Plot the predictions of internal and external contributions to  $B_r$  and  $B_\theta$  as a function of colatitude. How accurate are your results?