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RESEARCH ARTICLE

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Key Points:

- We analyze the fault length-frequency distribution in developing (structurally immature) fault systems
- The cumulative frequency distribution follows a power law over a range of fault length spanning 8 orders of magnitude, with a negative exponent of ∼2, consistent with the Gutenberg‐ Richter law
- Small faults within the brittle upper crust can accommodate a substantial (*>*30%) fraction of tectonic strain

Supporting Information:

Supporting Information may be found in the online version of this article.

Correspondence to:

X. Zou, x3zou@ucsd.edu

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Author Contributions:

Conceptualization: Yuri Fialko **Data curation:** Xiaoyu Zou **Formal analysis:** Xiaoyu Zou, Yuri Fialko **Funding acquisition:** Yuri Fialko **Investigation:** Xiaoyu Zou, Yuri Fialko **Methodology:** Xiaoyu Zou, Yuri Fialko **Project administration:** Yuri Fialko **Resources:** Yuri Fialko **Software:** Xiaoyu Zou, Yuri Fialko **Supervision:** Yuri Fialko **Validation:** Xiaoyu Zou, Yuri Fialko **Visualization:** Xiaoyu Zou, Yuri Fialko **Writing – original draft:** Xiaoyu Zou **Writing – review & editing:** Xiaoyu Zou, Yuri Fialko

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Can Large Strains Be Accommodated by Small Faults: "Brittle Flow of Rocks" Revised

Xiaoyu Zou¹ and Yuri Fialko¹

¹Institute of Geophysics and Planetary Physics, Scripps Institution of Oceanography, University of California San Diego, La Jolla, CA, USA

Abstract Brittle deformation in the upper crust is thought to occur primarily via faulting. The fault length‐ frequency distribution determines how much deformation is accommodated by numerous small faults versus a few large ones. To evaluate the amount of deformation due to small faults, we analyze the fault length distribution using high-quality fault maps spanning a wide range of spatial scales from a laboratory sample to an outcrop to a tectonic domain. We find that the cumulative fault length distribution is well approximated by a power law with a negative exponent close to 2. This is in agreement with the earthquake magnitude‐frequency distribution (the Gutenberg‐Richter law with b‐value of 1), at least for faults smaller than the thickness of the seismogenic zone. It follows that faulting is a self-similar process, and a substantial fraction of tectonic strain can be accommodated by faults that don't cut through the entire seismogenic zone, consistent with inferences of "hidden strain" from natural and laboratory observations. A continued accumulation of tectonic strain may eventually result in a transition from distributed fault networks to localized mature faults.

Plain Language Summary The Earth's crust is pervasively damaged, and contains faults, fractures, and joints of various sizes and orientations. We use mapped fault traces from multiple data sets spanning a wide range of scales to investigate how much deformation is accommodated by small versus large faults. The fault length distribution is often assumed to be fractal, that is, following a power law. The power-law exponent α quantifies the relative contributions of many small faults relative to a few large ones. For $\alpha \leq 1$, the contribution of small faults is negligible, while for $\alpha \geq 2$, strains accommodated by small faults become significant. We find that the cumulative fault length distribution approximately follows a power law with an exponent $\alpha \geq 2$. This implies that small faults in developing shear zones can accommodate a substantial fraction of tectonic strain.

1. Introduction

Tectonic deformation in the upper crust is mainly accommodated by brittle failure, manifested in faults and tensile cracks (e.g., S. Cox & Scholz, [1988](#page-12-0)). Faults are ubiquitous in both intraplate settings and at plate boundaries (e.g., Woodcock, [1986](#page-15-0); Bezerra et al., [2011](#page-11-0); R. T. Cox et al., [2001;](#page-12-0) Twiss & Moores, [1992;](#page-14-0) Bürgmann & Pollard, [1994\)](#page-12-0). As faults continue to slip, they increase their length via crack tip propagation, linkage, and coalescence (e.g., Mansfield & Cartwright, [2001;](#page-13-0) S. Cox & Scholz, [1988;](#page-12-0) Dawers & Anders, [1995;](#page-12-0) Fossen, [2020](#page-12-0); Rotevatn et al., [2019\)](#page-14-0). As a result, the upper crust contains faults of various sizes, from millimeter-long microfractures to mature faults extending hundreds of kilometers. The fault length distribution controls the relative contributions of small versus large faults to a total strain budget and is of interest to many disciplines including tectonics, engineering geology, hydrogeology, petroleum industry, and seismic hazards assessment (e.g., C. H. Scholz & Cowie, [1990](#page-14-0); Bense et al., [2013](#page-11-0); Bonnet et al., [2001](#page-12-0); Kolawole et al., [2019](#page-13-0)).

Previous studies suggested a variety of functional forms describing the fault size distribution. It is generally believed that in a low-strain environment (e.g., developing shear zones), fault populations are fractal and thus follow a power‐law distribution (e.g., Childs et al., [1990;](#page-12-0) D. Turcotte, [1986](#page-14-0); Bour & Davy, [1999;](#page-12-0) Bonnet et al., [2001](#page-12-0); Ben‐Zion & Sammis, [2003](#page-11-0)). Nicol et al. [\(1996](#page-13-0)) noted that the fault length distribution may deviate from a power-law if a wide range of fault lengths is considered, and that the power-law exponent may vary at the low end of the fault length distribution owing to spatial clustering. In contrast, Odling et al. [\(1999](#page-13-0)) argued that the fault length distribution may appear as log-normal in individual data sets with a given detection threshold, but is a power‐law for "composite" data sets that combine a number of individual data sets spanning a wide range of spatial scales. Other considered functional forms include gamma and exponential distributions that may provide a

better fit to the data, especially at the distribution tails (e.g., Ackermann et al., [2001](#page-11-0); Cowie et al., [1993](#page-12-0); Michas et al., [2015](#page-13-0); Spyropoulos et al., [1999\)](#page-14-0). However, it is not always clear if departures from a power law are real, or due to sampling artifacts (e.g., Bonnet et al., [2001](#page-12-0)). Gupta and Scholz ([2000\)](#page-12-0) suggested a transition from a power‐law to an exponential distribution when tectonic strain exceeds a critical threshold of the order of 0.1.

In case of a power‐law distribution, the number of faults *N* that have lengths greater than or equal to *L* is given by

$$
N(L) = CL^{-\alpha} \tag{1}
$$

where *C* is an empirical constant, and $\alpha > 0$ is an absolute value of the power-law exponent, also known as the Pareto index (e.g., Clark et al., [1999\)](#page-12-0). The derivative of the cumulative fault length distribution (Equation 1) with respect to *L* is the probability density,

$$
\frac{dN}{dL} = C(1 - \beta)L^{-\beta},\tag{2}
$$

which is also a power law, with $\beta = \alpha + 1$. The probability density (Equation 2) is sometimes referred to as the non‐cumulative frequency distribution. A number of studies used field observations to test the assumption of a fractal distribution, and estimate parameters *C* and *α* (or *β*). Reported values of the best‐fit power‐law exponent *α* vary from 0.7 for faults in Chimney Rock, Utah (Cladouhos & Marrett, [1996;](#page-12-0) Krantz, [1988\)](#page-13-0) to 1.1 for Neogene faults in the Boso and Iura Peninsula, Japan (C. H. Scholz & Cowie, [1990\)](#page-14-0) to 2.3 for faults and fractures in sandstone in Tayma, Saudi Arabia (Odling et al., [1999\)](#page-13-0). Most of the previous studies used data sets consisting of $10^2 - 10^3$ fault traces with fault lengths spanning 1–2 decades.

The magnitude of the power-law exponent determines how deformation is partitioned between small and large faults. C. H. Scholz and Cowie [\(1990](#page-14-0)) estimated the power-law exponent $\alpha \approx 1$ using fault trace data from Japan and concluded that small faults are negligible in the total strain budget. In contrast, Kautz and Sclater ([1988\)](#page-13-0) argued, based on laboratory experiments and observations of natural faults, that small-scale faulting is responsible for a substantial internal deformation within crustal blocks bounded by major faults. Such deformation was also invoked to explain the relative rotation of conjugate faults in tectonically active regions such as the Eastern California Shear Zone (Fialko & Jin, [2021\)](#page-12-0) and Tibet (Yin & Taylor, [2011](#page-15-0)). No such rotation would be possible if small faults were too scarce to accommodate a substantial fraction of tectonic strain.

To quantify the amount of deformation that can be attributed to small‐scale faulting, we analyze the fault length distribution across a wide range of spatial scales using several high‐quality data sets. In particular, we use detailed fault maps from different geological settings, including the Basin and Range Province (Nevada), Central Pennsylvania/Northern New Jersey, Ventura County (California), and Northern New Zealand. We complement these crustal-scale data sets with outcrop-scale observations of fractures and joints in Eastern Israel (Bahat, [1987\)](#page-11-0) and Sierra Nevada (Segall & Pollard, [1983\)](#page-14-0), and dip‐slip faults in Southern New Zealand (Davis et al., [2005\)](#page-12-0) and Eastern France (Villemin et al., [1995](#page-14-0)). We also use laboratory observations of Mode I and II micro-fractures in rock samples loaded to failure at confining pressures of several tens of megapascals (Katz & Reches, [2004\)](#page-13-0). We examine the compiled multi-scale data to test the assumption of a power-law distribution, obtain the best-fit power‐law exponent, and use the latter to estimate the amount of strain accommodated by faults in the upper crust, as a function of fault size.

2. Data and Methods

Continental deformation often involves broad regions of distributed faulting such as the Eastern California Shear Zone (Dokka & Travis, [1990;](#page-12-0) Floyd et al., [2020](#page-12-0); Tymofyeyeva & Fialko, [2015\)](#page-14-0), Basin and Range (Eaton, [1982](#page-12-0); Hodges et al., [1989\)](#page-13-0), and India‐Arabia‐Africa‐Eurasia collision zone (e.g., England, [1987](#page-12-0); Reilinger et al., [2006\)](#page-14-0).

We are interested in the fault length-frequency distribution in regions of distributed deformation. Unfortunately, strike‐slip faults are often difficult to recognize due to their limited geomorphologic expression, especially in case of small, low offset faults. Normal faults are better suited for this purpose because they produce scarps(changes in topography) that are easier to map. One of the most extensive and detailed fault trace data sets from an actively deforming extensional region is that from the Basin and Range (B&R) province in the Western US (Figure [1a\)](#page-2-0). This region hosts a number of active Quaternary faults (e.g., Eaton, [1982](#page-12-0); U.S. Geological Survey and Nevada

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Figure 1. (a) Map of the Basin and Range Province (U.S. Geological Survey and Nevada Bureau of Mines and Geology, [2023](#page-14-0)). Shading denotes topography. Black lines denote fault traces. Red lines denote fault traces excluded from the analysis as they intersect the region boundaries. Inset shows location of the area of interest (white rectangle) in a regional context; thin black lines indicate state boundaries. The concatenated fault data set includes 10,825 fault segments. The minimum segment length is 2.1 m and the maximum length is 49 km. (b) Probability density of the fault length distribution, on a log‐log scale. Solid line represents the best linear fit at the high end of the fault length distribution (*L >* 5 km). The estimated power-law exponent (slope of the best-fit line) is $\beta = 3.51 \pm 0.12$.

Bureau of Mines and Geology, [2023\)](#page-14-0). We examine fault traces from an area extending 6° in longitude and 4° in latitude (Figure 1a). The respective data set consists of 26,512 fault traces, with the fault segment lengths varying from 2.1 m to 42.6 km.

A close inspection of the B&R fault trace data reveals that many fault traces that appear continuous on a regional scale (Figure 1a) are in fact highly segmented (Supplementary Figure S1a in Supporting Information S1). While some of the apparently continuous fault traces may be segmented because they have different attributes such as dip and strike, others may have the same attributes but are still separated at the segment level. To mitigate potential biases due to artificial segmentation, we developed an algorithm for concatenating individual segments that likely belong to the same fault. The algorithm attributes different segments to the same fault if the following criteria are satisfied: (a) tips of the adjacent fault segments are within a prescribed distance *D* from each other; (b) the adjacent fault segments are sufficiently well aligned, such that the difference in strike angles θ_1 and θ_2 between the segment tips (see Figure 2a) is less than a prescribed threshold *δ*; also, the difference between the average of strike angles at the segment tips, $(\theta_1 + \theta_2)/2$ and the strike angle of a line connecting the segment tips

Figure 2. A schematic illustrating criteria used for connecting fault segments. Black solid lines denote fault traces (segments f1 and f2). *D* is the distance between fault tips. θ_1 and θ_2 are the local strike angles at the fault tips. The difference in strike angles is $\delta = |\theta_1 - \theta_2|$. (a) Segments f1 and f2 are allowed to be connected if the fault traces are sufficiently close and aligned (*D* and δ are below the prescribed thresholds). (b) Example of a configuration when the fault tips are not aligned (δ is larger than the prescribed threshold). (c) Example of a configuration of sub-parallel faults. In our analysis, we use $D \le 5$ km and $\delta \le 30^\circ$.

is less than a prescribed threshold *δ* (Figure [2b\)](#page-2-0); (c) overlapping segments that satisfy conditions (a) and (b) are considered part of the same fault if $D < L/3$, where L is the length of a smaller segment. The latter condition is meant to avoid absorption of small faults that are sub-parallel to (rather than aligned with) the large ones (Figure [2c\)](#page-2-0). The respective criteria are illustrated in Figure [2](#page-2-0).

A reasonable upper limit on *D* is some fraction of the thickness of the brittle layer *T*, such that the apparently discontinuous (e.g., poorly exposed) surface traces might possibly belong to the same fault at depth. For the Basin and Range province, $T \approx 15$ km (e.g., Pancha et al., [2006](#page-14-0)). We assume $D \le 5$ km. This assumption is consistent with observations and models of earthquake ruptures jumping across nearby fault segments (e.g., Ando & Kaneko, [2018;](#page-11-0) Harris & Day, [1999](#page-13-0); Jia et al., [2023](#page-13-0)). We find that the best-fit power-law exponent is relatively insensitive to the assumed value of *D*, for *δ* between 0 and 30° (Figures S2 and S3 in Supporting Information S1). Larger values of *D* and δ encourage segment linking, resulting in a smaller number of small faults, and consequently smaller absolute values of the best-fit power-law exponents. In our analysis, we use $D = 5$ km, and $δ = 30°$ to provide a lower bound on *α*. A comparison of fault trace data before and after "de-segmentation" is shown in Figure S1 in Supporting Information S1.

Because the cumulative fault length distribution is known to be sensitive to finite size effects, which can bias determination of the exponent (e.g., Bonnet et al., [2001](#page-12-0); Serafino et al., [2021\)](#page-14-0), we use the density distribution (Equation [2\)](#page-1-0) to estimate the power-law exponent β , unless indicated otherwise. The respective values of α are trivially given by $\alpha = \beta - 1$.

Figure [1b](#page-2-0) shows the probability density of fault length distribution for the "concatenated" Basin and Range data set (a subset is shown in Figure S1b in Supporting Information S1). To minimize the censoring bias (e.g., Torabi & Berg, [2011\)](#page-14-0), we refine the data set by excluding faults that intersect the region boundaries (after the segment concatenation), see red lines in Figure [1a](#page-2-0). On a log-log plot, the density distribution exhibits a quasi-linear trend for *L >* 5 km, and flattens out for smaller *L*. The roll‐off at *L <* 5 km likely results from incomplete sampling (truncation bias, Torabi & Berg, [2011;](#page-14-0) Bonnet et al., [2001](#page-12-0)), analogous to saturation of the Gutenberg‐Richter distribution below the magnitude of completeness (e.g., Woessner & Wiemer, [2005](#page-15-0)). The truncation bias may be due to a finite detection threshold and/or 2-D sampling of a 3-D fault population (e.g., Heifer & Bevan, [1990](#page-13-0); Marrett & Allmendinger, [1991\)](#page-13-0). We use the Kolmogorov-Smirnov (KS) test (Clauset et al., [2009](#page-12-0)) to identify the range of fault lengths $[L_{min}, L_{max}]$ that can be used for power-law fitting (see Supplementary Text S1 in Supporting Information S1 for details). We estimate the density power-law exponent β by the least squares linear regression over the interval [*L*_{min}, *L*_{max}]. The uncertainty on the best-fit slope is obtained by performing a regression for different bin sizes, and computing a standard deviation of the resulting slope estimates. For the data shown in Figure [1](#page-2-0), we obtain $\beta = 3.51 \pm 0.12$, or $\alpha \approx 2.5$. This can be compared to the value of $\alpha = 1.84$ estimated by Cladouhos and Marrett [\(1996](#page-12-0)), who used an older (presumably, less complete) fault map of the Basin and Range province, and fitted a linear trend to the cumulative fault length distribution over the fault length interval between ∼15–70 km.

We extended the same analysis to several other locations for which high-resolution maps of dip-slip faults are openly available, in particular, Central Pennsylvania and Northern New Jersey, Ventura County (California), and Northern New Zealand. Figure [3a](#page-4-0) shows fault traces from an area in Central Pennsylvania and Northern New Jersey (NJ Dept. of Environmental Protection Bureau of GIS, [2023;](#page-13-0) PA Department of Conservation & Natural Resources, [2023](#page-13-0)). The mapped traces represent inactive thrust and strike‐slip faults formed 400 to 250 million years ago (Hatcher, [1987](#page-13-0)). For consistency, we apply the same algorithm for concatenating the aligned segments as described above. The resulting data set consists of 2,273 faults having length between 15 m and 108 km. The probability density fault length distribution (Figure [3b](#page-4-0)) is characterized by an apparent truncation for faults smaller than 20 km, and a slope of the quasi-linear trend of -3.51 , remarkably similar to results obtained for the Basin and Range province (Figure [1b](#page-2-0)).

The Ventura County, CA (Figure [4](#page-4-0)) and Northern New Zealand (Figure [5](#page-5-0)) fault maps cover much smaller areas. After the segment concatenation procedure, each data set contains several hundreds of fault traces. This is 1–2 orders of magnitude smaller than the number of fault traces in the B&R and Pennsylvania/New Jersey data sets (Figures [1](#page-2-0) and [3](#page-4-0)), but comparable to a typical size of data sets examined in a number of previous studies. While these smaller data sets are too characterized by decaying trends toward the high end of the sampled range of fault lengths, the data exhibit a significant scatter (e.g., Figure [5b](#page-5-0)), making power-law fits more problematic. Our

Figure 3. (a) Map of Central Pennsylvania and Northern New Jersey (NJ Dept. of Environmental Protection Bureau of GIS, [2023](#page-13-0); PA Department of Conservation & Natural Resources, [2023\)](#page-13-0). Notation is the same as in Figure [1.](#page-2-0) The concatenated fault data set includes 2,273 fault segments. The minimum segment length is 15 m and the maximum length is 108 km. (b) Probability density of the fault length distribution, on a log‐log scale. Solid line representsthe best linear fit at the high end of the fault length distribution (*L* > 10 km). The estimated power-law exponent (slope of the best-fit line) is $\beta = 3.51 \pm 0.20$.

analysis of the respective data sets yields smaller values of *β* that are subject to higher uncertainties (2*.*68 ± 0*.*14 for Ventura County and 2*.*42 ± 0*.*40 for Northern New Zealand, see Figures 4b and [5b](#page-5-0)).

To evaluate the fault length distribution at smaller scales, we use published data on fracture density measured in outcrops (*L* ∼ 1–100 m) and laboratory samples (*L* ∼1–100 mm). The outcrop‐scale observations include reactivated joints in igneous rocks near Florance Lake, Sierra Nevada, California (Segall & Pollard, [1983\)](#page-14-0) and Eocene chalks in the Syrian Arc folding belt, Israel (Bahat, [1987\)](#page-11-0); thrust faults in the Ostler Fault Zone, Benmore outcrop, Southern New Zealand (Davis et al., [2005](#page-12-0)); and predominantly dip-slip faults in La Houve Coal Field, an old sedimentary basin in Eastern France that experienced both compressional and extensional tectonics (Villemin et al., [1995\)](#page-14-0). The laboratory data are from specimens of Mount Scott granite of Oklahoma loaded to peak yield stress in a triaxial apparatus under confining pressure of 41 MPa (Katz & Reches, [2004](#page-13-0)). The micro-structural mapping of the sample damage was performed on scanned images of thin sections. Each sample had on the order of 10^3 resolved micro-fractures with lengths between 0.01 and 10 mm (Katz & Reches, [2004](#page-13-0)).

A compilation of the respective data sets is presented in Figure [6](#page-5-0), along with the fault trace data from Figures [1](#page-2-0) and 3–5. To enable a direct comparison of different data sets, we normalize the cumulative fault length counts by the areas from which the fault trace data were collected. The combined cumulative frequency distribution spans

Figure 4. (a) Map of Ventura County, CA (County of Ventura, [2023\)](#page-12-0). Notation is the same as in Figure [1](#page-2-0). The concatenated fault data set includes 349 fault segments. The minimum segment length is 0.6 m and the maximum length is 30 km. (b) Probability density of the fault length distribution, on a log‐log scale. Solid line represents the best linear fit at the high end of the fault length distribution $(L > 3 \text{ km})$. The estimated power-law exponent (slope of the best-fit line) is $\beta = 2.68 \pm 0.14$.

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Figure 5. (a) Map of Northern New Zealand (Langridge et al., [2016a\)](#page-13-0). Notation is the same as in Figure [1](#page-2-0). The concatenated fault data set includes 159 fault segments. The minimum segment length is 363 m and the maximum length is 24.7 km. (b) Probability density of the fault length distribution, on a log-log scale. The solid line represents the best linear fit at the high end of the fault length distribution $(L > 4 \text{ km})$. The estimated power-law exponent (slope of the best-fit line) is $\beta = 2.42 \pm 0.40$.

8 decades of fault length, and 18 decades of fault density (cumulative fault counts per unit area). All of the individual data sets shown in Figure 6 appear to have a log-normal distribution, with a quasi-linear trend at the high end, and a roll-off at the low end of the respective fault lengths. The high end quasi-linear trends have slopes that

Figure 6. Cumulative fault length frequency distribution for a combined data set including fault traces (Figures [1–4](#page-2-0)), as well as outcrop-scale and lab data, normalized by the respective observation areas, on a log-log scale. The black solid line is the least squares fit to the "high‐end" asymptotes of the outcrop‐ and crustal‐scale data (black symbols, upper right legend). The estimated power-law exponent is $\alpha = 2.06$, and the pre-multiplying factor is $C = 0.53$. Blue triangles denote micro-crack length distributions from laboratory tests (Katz & Reches, [2004\)](#page-13-0). Black triangles denote the same data shifted toward the best-fit envelope (black solid line).

are quite similar for all data sets shown in Figure [6](#page-5-0), indicative of similar power-law exponents. Furthermore, the crustal‐ and outcrop‐scale data admit a common envelope, suggesting that not only the power‐law exponent, but also the multiplier *C* (Equation [1\)](#page-1-0) is essentially the same across multiple data sets. The least squares fit of the common envelope to the crustal- and outcrop-scale data (see black solid line in Figure [6\)](#page-5-0) yields a power-law exponent of $\alpha = 2.06$ $\alpha = 2.06$. The micro-crack data (blue triangles in Figure 6) parallel this trend, but fall somewhat below, suggesting a lower value of *C* (i.e., a smaller crack density). The least squares fit of an envelope to all of the data, including the micro-crack data (blue triangles in Figure [6](#page-5-0)), yields a power-law exponent of $\alpha = 1.93$ (not shown). To illustrate the overall similarity of the slope of the micro-crack length distribution to the rest of the data, we also plot the micro-crack data shifted toward the best-fit envelope ($\alpha = 2.06$, see black triangles in Figure [6\)](#page-5-0).

3. Strain Due To Faults Obeying a Power Law Distribution

An overall agreement of the estimated power-law exponents of individual data sets between each other, on the one hand, and the common envelope, on the other hand (Figure [6](#page-5-0)), lends support to a suggestion that the roll-off in individual data sets is a result of truncation (e.g., due to a detection threshold, Bonnet et al., [2001](#page-12-0); Torabi $\&$ Berg, [2011\)](#page-14-0), and that the fault length statistics is adequately described by a power law across a wide range of spatial scales. If so, one can evaluate the amount of tectonic strain absorbed by faults of different sizes (e.g., C. H. Scholz & Cowie, [1990](#page-14-0); J. Walsh et al., [1991](#page-14-0)).

For a population of *n* faults within the brittle crust having a volume *TA*, where *T* is the thickness of the brittle layer, and *A* is the map area, the average strain accommodated by faulting is given by (Kostrov, [1974](#page-13-0)):

$$
\varepsilon_{ij} = \frac{1}{2TA} \sum_{k=1}^{n} {^k}P_{ij}.
$$
 (3)

In Equation 3, ${}^kP_{ij}$ is the seismic potency tensor (e.g., Ben-Zion, [2001\)](#page-11-0) of the *k*-th fault in a population. The average fault slip *S* is expected to scale with fault length *L*,

$$
S \propto L^m. \tag{4}
$$

Theoretical arguments and field observations suggest that *m* should be close to one (e.g., Cowie & Scholz, [1992](#page-12-0); Fialko, [2015](#page-12-0)), although higher values of *m* were suggested as well (e.g., J. J. Walsh & Watterson, [1988;](#page-14-0) Marrett & Allmendinger, [1991\)](#page-13-0). Assuming $m = 1$,

$$
S = \epsilon L,\tag{5}
$$

where ϵ is the critical shear strain drop corresponding to fault propagation. The scalar potency is $P = \gamma S L^2$ for faults smaller than *T*, and $P = \gamma SLT$ otherwise, where γ is a geometric factor of the order of unity that accounts for the fault shape and fault dip (for faults that cut through the entire brittle layer, e.g., Vavra et al., [2023](#page-14-0)). For simplicity, hereafter we assume $\gamma = 1$. The number of faults within an interval of fault lengths ΔL is (*dN*(*L*)*/dL*)Δ*L*. The cumulative potency can be calculated by integrating potencies of all faults for a given range of fault lengths. For faults smaller than *T*, the cumulative potency is (C. H. Scholz & Cowie, [1990\)](#page-14-0):

$$
p_1(L_{\min}, L_{\max}) = \sum_{k} {}^{k}P = -\epsilon \int_{L_{\min}}^{L_{\max}} \frac{dN(L)}{dL} L^3 dL = C \epsilon \frac{\alpha}{3 - \alpha} L^{3 - \alpha} \Big|_{L_{\min}}^{L_{\max}}, \tag{6}
$$

where L_{min} and L_{max} are the minimum and maximum fault sizes, respectively. For faults that cut through the entire brittle layer $(L > T)$,

$$
p_2(L_{\min}, L_{\max}) = \sum_{k} {}^{k}P = -\epsilon T \int_{L_{\min}}^{L_{\max}} \frac{dN(L)}{dL} L^2 dL = C\epsilon T \frac{\alpha}{2 - \alpha} L^{2 - \alpha} \Big|_{L_{\min}}^{L_{\max}}.
$$
 (7)

We evaluate the relative contribution of faults smaller than a given size *L* to the total strain by allowing $L_{\text{min}} \to 0$, and computing a ratio

Figure 7. Percentage of the total potency R (Equation 8) accommodated by faults having length less than *L*, for several estimated values of the powerlaw exponent *α*: solid line, $α = 2.06$ (this study); dotted line, $α = 2.34$ (Odling et al., 1999); dashed line, $\alpha = 1.1$ (C. H. Scholz & Cowie, [1990\)](#page-14-0). We assume L_{max} = 100 km (Figure [6](#page-5-0)). The vertical gray line corresponds to $L = T = 15$ km.

$$
R = 100\% \times \begin{cases} \frac{p_1(0, L)}{p_1(0, T) + p_2(T, L_{\text{max}})}, & \text{for } L < T\\ \frac{p_1(0, T) + p_2(T, L)}{p_1(0, T) + p_2(T, L_{\text{max}})}, & \text{for } L > T. \end{cases} \tag{8}
$$

Note that *R* does not depend on factors *C* and ϵ . Figure 7 shows the percentage of strain accommodated by faults having length less than *L*, for a range of *L*, assuming $\alpha = 2.06$, $L_{\text{max}} = 100$ km (Figure [6\)](#page-5-0), and $T = 15$ km, typical of the seismogenic depth in many tectonically active areas (e.g., Pancha et al., [2006;](#page-14-0) E. O. Lindsey & Fialko, [2016](#page-13-0); Jin et al., [2023;](#page-13-0) Jia et al., [2023\)](#page-13-0). For a comparison, we also show analogous calculations for previously reported values of $\alpha = 1.1$ (dashed line, C. H. Scholz & Cowie, [1990](#page-14-0)) and $\alpha = 2.34$ (dotted line, Odling et al., [1999](#page-13-0)).

4. Discussion

For fault systems characterized by a power‐law size distribution (1), the power-law exponent α controls how much of tectonic deformation is accommodated by numerous small faults versus a few large ones. C. H. Scholz and Cowie [\(1990](#page-14-0)) estimated the value of $\alpha = 1.1$ for a set of intraplate faults in Japan, and concluded that small faults are negligible in the overall strain budget. This is because integrals (6) and (7) are strongly convergent for $\alpha \approx 1$, so that the cumulative potency is dominated by the largest faults. Our

results, based on a much larger data set, indicate $\alpha \geq 2$ (Figure [6\)](#page-5-0). Most of the previously published estimates of *α* fall in the range between 1 and 2 (e.g., Bonnet et al., [2001](#page-12-0)). Possible reasons for different values of *α* reported in the literature include: (a) use of fault trace data of limited coverage and/or resolution; (b) uncertainties involved in defining fault connectivity; (c) a narrow range of fault lengths used in the analysis; (d) departures from selfsimilarity due to the presence of intrinsic length scales; (e) different stages of maturity of different fault systems. For example, the data set used by C. H. Scholz and Cowie [\(1990](#page-14-0)) spans only one order of magnitude of fault lengths, from ∼10 to ∼100 km, likely insufficient for a robust validation of a power-law distribution (Stumpf & Porter, [2012](#page-14-0)). C. Scholz et al. ([1993](#page-14-0)) analyzed a data set from the Volcanic Tableland (California) with fault lengths spanning 2 orders of magnitude, from a few tens of meters to a few kilometers, and obtained a higher value of $\alpha \approx 1.3$. The latter under-predicts the slope at the upper tail of the fault length distribution of C. Scholz et al. ([1993,](#page-14-0) their Figure 4), which the authors attributed to data censoring.

Our analysis of several high-resolution data sets (Figures 1, [3–6](#page-2-0)) suggests values of α close to 2, higher than those reported by C. H. Scholz and Cowie [\(1990](#page-14-0)) and C. Scholz et al. [\(1993](#page-14-0)), but consistent with results from other multi-resolution studies. In particular, Heifer and Bevan ([1990\)](#page-13-0) combined fault trace data with measurements of crack density in boreholes to infer $\alpha \approx 2$. Odling et al. [\(1999](#page-13-0)) performed a multi-scale analysis of the length distribution of faults and joints in sandstones in Saudi Arabia, and found the best-fit power-law exponent of 2.34 for a range of fault lengths spanning 4 orders of magnitude. Values of $\alpha \geq 2$ in sandstones may be due to strainhardening deformation bands that may inhibit fault propagation and instead promote nucleation of numerous small faults. C. Scholz et al. ([1993\)](#page-14-0) cautioned against combining observations that include different fracture modes (e.g., faults and joints, Heifer & Bevan, [1990\)](#page-13-0). However, it can be argued that the crack length distributions should not strongly depend on the fracture mode as mathematical expressions for stress fields due to shear and tensile cracks are essentially identical (e.g., Fialko, [2015](#page-12-0)), so that stress interactions within the crack network are expected to be similar (e.g., for shear and tensile cracks). This is consistent with results from previous studies. As noted by Bonnet et al. [\(2001](#page-12-0)), "…scaling exponents (notably, the length distribution exponent) are remarkably insensitive to the orientation of the slip vector, that is, whether or not fracturing is accompanied by shear (faults) or tensile (joints) displacement." Given that the total accommodated strain includes contributionsfrom all types of fractures, it's not unreasonable to include fractures of different modes in the length‐frequency analysis. Different data sets are expected to show some variability in the estimated power-law exponents given various lithologies, crustal layering, maturity states, and fault roughness, among other factors (e.g., Kim & Sanderson, [2005](#page-13-0); Power et al., [1988](#page-14-0)). Nevertheless, an overall agreement between the estimated power-law exponents for different types

Figure 8. A cumulative fault length‐frequency distribution of earthquakes that occurred in Southern California in 1981–2023. Open circles: earthquake data from the waveform-relocated catalog of Hauksson et al. ([2012,](#page-13-0) updated annually), $1.7 \leq M \leq 6.5$. Solid gray line: best linear fit in log-log coordinates. Slope of the best‐fit line is equal to − 1.998.

of fractures, as well as for data sets from different locations (Figure [6](#page-5-0)) lends support to a hypothesis that early stages of faulting may be governed by a "universal" power law with $\alpha \approx 2$ (King, [1983](#page-13-0); Proekt et al., [2012](#page-14-0); Roman & Bertolotti, [2022](#page-14-0)).

The relative contribution of small faults to the strain budget is expected to be larger for smaller values of L_{max} , and/or larger values of α . We note that the actual values of α may be in fact higher than those estimated from twodimensional (2‐D) sampling of three‐dimensional (3‐D) fault populations. For example, for uniformly distributed and randomly oriented faults, the true (i.e., 3‐D) exponent is predicted to be larger than the exponent inferred from the 2-D sampling by as much as 1 unit (e.g., Bonnet et al., 2001 ; Marrett & Allmendinger, [1991\)](#page-13-0). This only applies to small $(L < D)$ faults, as for large faults the distribution is essentially 2-D. For α approaching 3, small faults would actually dominate the strain budget, and the contribution of large faults would be negligible. Note that for the cumulative potency and strain to remain finite, α cannot exceed 3 (Equation [6](#page-6-0)).

Some useful constraints on the fault length‐frequency distribution can be obtained from the magnitude‐frequency distribution of earthquakes. The latter obeys the well-established Gutenberg-Richter law, $log_{10} N(M) = a - bM$, where $N(M)$ is the number of earthquakes with magnitude larger or equal to M , and a and b are empirical constants

(Gutenberg & Richter, [1944\)](#page-12-0). Unless otherwise noted, *M* denotes the moment magnitude, $M = (\log_{10} M_0 - 9.05)/1.5$, where M_0 is the scalar seismic moment (Hanks & Kanamori, [1979\)](#page-13-0). Because *M* is related to a logarithm of M_0 , it can be shown that parameter *b* (referred to as the "b-value") defines a power-law exponent *αs* in the size‐frequency distribution of seismic ruptures (e.g., King, [1983;](#page-13-0) D. L. Turcotte, [1997](#page-14-0)). In particular, the scalar seismic moment is given by $M_0 = \mu AS$, where μ is the shear modulus, A is the rupture area, and *S* is the average coseismic slip. For ruptures with moderate aspect ratios, $A \propto L^2$, and $S \propto L$, so that $M_0 \propto L^3$ (e.g., Fialko, [2015](#page-12-0); Kanamori & Anderson, [1975](#page-13-0)). Substituting this scaling relation into expressions for the moment magnitude and magnitude-frequency distribution, one obtains $N(L) \propto L^{-2b}$, or $\alpha_s = 2b$. The coefficient of proportionality *C* can be found under additional assumptions, for example, by approximating earthquake ruptures as circular cracks with a constant stress drop $\Delta \sigma$. In this case, $A = \pi L^2/4$, and *S* = 8Δ*σL*/7*πμ* (e.g., Eshelby, [1957](#page-12-0)), leading to $L = (7M_0/2Δσ)^{1/3}$. Figure 8 shows a cumulative sizefrequency distribution of earthquakes in Southern California recorded over the last 42 years since the deploy-ment of digital seismic networks (Hauksson et al., [2012](#page-13-0)). The data set includes 7.7×10^5 earthquakes; shown in Figure 8 are events in the magnitude range from 1.7 (magnitude of completeness) to 6.5 (saturation of the seismogenic layer, $L \leq T$). To estimate rupture sizes from the moment magnitude, we choose a constant stress drop of 3 MPa (Abercrombie, [1995;](#page-11-0) Shearer et al., [2006](#page-14-0)). Note that the assumed value of Δ*σ* does not affect the power‐law exponent *αs* given stress drops do not systematically depend on *L* (Allmann & Shearer, [2009\)](#page-11-0).

The earthquake data unambiguously indicate that seismically active faults within the seismogenic layer ($L < T$) have a power-law length-frequency distribution with an exponent α_s close to 2, corresponding to a b-value close to 1. The same b-value well describes the global seismicity, with some variations between 0.7 and 1.4 depending on the area and tectonic regime (Godano & Pingue, [2000](#page-12-0); Schorlemmer et al., [2005](#page-14-0); Shearer et al., [2006](#page-14-0)). For earthquakes that rupture the entire seismogenic zone ($M > 6.5$ for $T = 12 - 15$ km in most of the Western North America), the seismic moment no longer scales as a cube of *L*, leading to values of α_s that are smaller than 2*b*. In particular, for $L > T$ the fault area is proportional to L , $A = LT$. The coseismic slip *S* is expected to saturate for $L > T$, but it appears to do so only for $L \gg T$ (e.g., Shaw & Scholz, [2001\)](#page-14-0), giving rise to empirical scaling of *S* in proportion to either *L* (C. Scholz, [1994\)](#page-14-0) or \sqrt{L} (Leonard, [2010](#page-13-0)). Correspondingly, the seismic moment scales as either L^2 or $L^{3/2}$, yielding values of α_s of 4*b*/3 or *b*. Assuming $b = 1$, this would suggest a change in the fault length-frequency distribution around $L \sim T$, with $\alpha_s \approx 2$ for "small" faults ($L < T$), and $\alpha_s \approx 1 - 1.3$ for "large" faults ($L > T$). However, it is unclear if b-value remains constant across the saturation length scale *L* ∼ *T*. Regional data sets such as that shown in Figure 8 contain too few events with magnitudes greater than 6.5 to robustly evaluate the length‐frequency distribution of earthquakes that rupture the entire

seismogenic zone. To gain further insight, we use the global CMT catalog (Ekström et al., [2012\)](#page-12-0) and examine the magnitude‐frequency distribution of strike‐slip earthquakes that occurred between 1976 and 2024 (see Figures S4 and S5 in Supporting Information \vert S1). In the magnitude range between 5 and 6.5, the estimated b-value of the global data set is about 1 (Figure S4 in Supporting Information S1), consistent with the more complete Southern California data set spanning almost 5 units of earthquake magnitude, $1.7 \leq M \leq 6.5$ (Figure [8\)](#page-8-0). However, strike-slip earthquakes that rupture the whole seismogenic zone, $6.5 \leq M \leq 8.1$, appear to have a larger b-value of ∼1.3 (Figure S4 in Supporting Information S1). This gives rise to *α_s* of 1.3 (assuming *S* ∝ √*L*) to 1.8 (assuming *S* ∝ *L*) for "large" faults ($L > T$).

An excellent agreement between results shown in Figures [6](#page-5-0) and [8](#page-8-0) may be interpreted as indicating that the lengthfrequency distribution of seismically active faults is the same as that of the entire fault population, $\alpha_s \approx \alpha$. Alternatively, it can be argued that α_s is a lower bound on α , as only a fraction of faults within the seismogenic zone is seismically active (e.g., D. L. Turcotte, [1997](#page-14-0), p. 44). This would be the case if estimates of α based on statistics of mapped fault traces suffer from a lower-dimensional bias (for $L < T$). The latter interpretation may be consistent with values of $\alpha > 2$ estimated from data sets that include large faults ($L > T$, Figures [1](#page-2-0) and [3\)](#page-4-0). Note that neither the fault length statistics of large $(L > T)$ faults derived from the fault trace data, nor the values of *αs* derived from the earthquake catalog data are subject to a lower‐dimensional bias. In other words, *α* can be larger than 2 for the entire fault length distribution.

In any case, it is clear that values of $\alpha \sim 1$ reported in some previous studies are inconsistent with the Gutenberg-Richter statistics at the low end of the fault size distribution. For example, C. H. Scholz and Cowie ([1990\)](#page-14-0) suggested $\alpha = 1.1$ based on analysis of a data set dominated by large faults ($L \sim 10 - 100$ km). While this result might be reconciled with seismic observations in the limit $L > T$, it fails to describe the length-frequency distribution of small faults for which $\alpha \geq 2$ (Figure [6\)](#page-5-0) as well as small-to-moderate earthquakes for which $\alpha_s \approx 2$ (Figure [8](#page-8-0)). In addition, our data sets that span the likely transition from "small" to "large" faults do not show obvious breaks in scaling and/or low values of *α* (Figures [1](#page-2-0) and [3\)](#page-4-0). However, the level of completeness of individual fault trace data sets is admittedly insufficient to make robust conclusions about potential variations in *α* across the presumed saturation length scale *L* = *T*.

We note that faults having lengths of tens to hundreds of kilometers have accumulated a substantial amount of slip, and thus may be more representative of a structurally mature fault system. Experimental studies reveal higher values of α at the initial stages of faulting when deformation is broadly distributed, and a decrease in α with an increasing system maturity (e.g., Cladouhos & Marrett, [1996](#page-12-0); Hatton et al., [1993](#page-13-0); Sornette et al., [1993](#page-14-0)). It follows that small faults can potentially accommodate a substantial fraction of tectonic strain at the initial stages of faulting (e.g., in developing shear zones). Over time, as faults grow and connect, deformation may localize to major faults that eventually take up most of the deformation.

These arguments suggest a distinction between deformation styles due to immature shear zones such as the Eastern California Shear Zone (Dokka & Travis, [1990](#page-12-0); Floyd et al., [2020\)](#page-12-0), and mature well‐slipped plate boundary faults such as the San Andreas Fault (Fialko, [2006;](#page-12-0) Lisowski et al., [1991\)](#page-13-0). In the latter case, interseismic strain accumulation is equal in magnitude, but opposite in sign to strain released in large earthquakes, so the patterns of interseismic and long‐term (geologic) displacements across a mature fault are very different (Figure [9\)](#page-10-0). A complete or nearly complete recovery of interseismic strain (i.e., elastic rebound) is evidenced by the good agreement between "geologic" and "geodetic" slip rates on major plate boundary faults (e.g., Schmalzle et al., [2006](#page-14-0); Tatar et al., [2012](#page-14-0); E. Lindsey & Fialko, [2013\)](#page-13-0). In contrast, immature fault systems with *α* ≥ 2 give rise to a distributed inelastic deformation with the long‐term displacement profile that may closely mimic the observed interseismic velocities (Fialko & Jin, [2021](#page-12-0)). The diffuse deformation pattern illustrated in Figure [9a](#page-10-0) can be thought of as resulting from the "seismic flow of rocks", as originally envisioned by Riznichenko ([1965\)](#page-14-0) and Kostrov ([1974\)](#page-13-0), although a more appropriate term would be the "brittle flow of rocks", since some of the deformation may occur aseismically, for example, via fault creep (Kaneko et al., [2013;](#page-13-0) Tymofyeyeva et al., [2019](#page-14-0); Vavra et al., [2024](#page-14-0)). Additionally, some inelastic strain can result from other deformation mechanisms such as folding, bulk plasticity, pressure solution, etc (Donath & Parker, [1964;](#page-12-0) Hamiel et al., [2006;](#page-12-0) Hancock, [1985](#page-12-0)). The "brittle strain" estimated from the fault length‐frequency statistics should thus be considered a lower bound on the total inelastic strain accommodated by the upper crust.

Figure 9. Schematic representation of kinematics of (a) developing shear zone and (b) mature plate boundary fault. Top and bottom panels denote interseismic and long-term (averaged over multiple earthquake cycles) motion, respectively. Gray lines denote active faults.

Taking at face value the estimated power law-exponent $\alpha \approx 2$ (Figure [6\)](#page-5-0), we find that small $(L < T)$ faults may take up more than one third of the total strain, which is much larger than predicted for $\alpha \approx 1$ (Figure [7](#page-7-0)). A power law-exponent $\alpha \geq 2$ may provide an explanation for the "missing strain" in palinspastic restorations of faults in sedimentary basins, as well as in laboratory models of tectonic extension using analog materials (e.g., Kautz & Sclater, [1988;](#page-13-0) Marrett & Allmendinger, [1992](#page-13-0); J. Walsh et al., [1991](#page-14-0)). The bulk inelastic deformation accommodated by small faults can result in rotation of faults away from the optimal orientation, and increasesin dihedral angles between conjugate faults, as often observed in active shear zones (e.g., Fialko, [2021](#page-12-0); Ron et al., [2001;](#page-14-0) Zou et al., [2023](#page-15-0)). It might also account for the reported differences between geologic and geodetic slip rates in regions of diffuse deformation. In particular, models of deformation across the plate boundary in California suggest that up to 30*%* of deformation is accommodated off of the known faults (Field et al., [2014](#page-12-0)). Similar conclusions are drawn from numerical models of continental extension (Pan et al., [2023](#page-14-0)). Given no resolvable difference between the geologic and geodetic slip rates of mature high-slip-rate faults such as the San Andreas and San Jacinto faults (Segall, [2002;](#page-14-0) E. O. Lindsey et al., [2014](#page-13-0); Tymofyeyeva & Fialko, [2018](#page-14-0); Schmalzle et al., [2006](#page-14-0)), most of the "missing slip" is apparently associated with regions of diffuse deformation characterized by low strain rates such as the Eastern California Shear Zone (Herbert et al., [2014\)](#page-13-0). The same may apply to other areas of broadly distributed continental deformation such as the India‐Eurasia collision zone (e.g., Garthwaite et al., [2013;](#page-12-0) Wang & Shen, [2020;](#page-15-0) Yin & Taylor, [2011](#page-15-0)). A major outstanding question is how strongly the "true" power‐law exponent α differs from estimates based on the fault trace data in case of small ($L < T$) faults, and/or "seismic" values of α_s based on the Gutenberg-Richter statistics. Values as high as 2.5 (or greater) would imply a diffuse long-term deformation similar to a viscous flow. Our results provide a lower bound on the amount of permanent strain accommodated in the upper crust off of major faults. Additional strain may be accommodated by other mechanisms such as pressure solution and bulk creep (e.g., Gratier et al., [2013](#page-12-0); Hancock, [1985\)](#page-12-0). Quantifying the respective contributions is a rich area for future research.

5. Conclusions

We analyzed the fault length frequency distribution using high-resolution fault trace data from diverse settings including Basin and Range Province, Central Pennsylvania/Northern New Jersey, Ventura County, California, and Northern New Zealand. To extend our analysis to smaller scales, we included published outcrop data from Sierra Nevada, Eastern Israel, Southern New Zealand, and Eastern France, and laboratory data from experiments on the initially intact granite samples. Our results indicate that while each individual data set yields an apparent log‐normal distribution of fault lengths, a composite multi‐scale data set reveals a fault length‐distribution that follows a power law over 8 decades of fault lengths, with a cumulative power-law exponent $\alpha \approx 2$. This is consistent with the Gutenberg-Richter statistics of earthquakes with a typical b-value of 1. However, not all faults present in the seismogenic zone may be seismically active, and the obtained value of α may under-estimate the true value of the power‐law exponent for example, due to an observation bias (2‐D sampling of 3‐D faults). We used the best-fit value of the power-law exponent to estimate the fraction of strain accommodated by faults as a function of fault size. We find that small faults (*L <* 15 km) can accommodate a substantial (more than 30%) fraction of tectonic strain, at least at the initial stages of faulting. This fraction may be substantially higher if the

fault length statistics suffer from a low-dimensional bias, and the true value of α is between 2 and 3. A continued deformation may give rise to a transition from distributed fault networks to highly localized mature faults, associated with a decrease in *α*.

Conflict of Interest

The authors declare no conflicts of interest relevant to this study.

Data Availability Statement

The fault and fracture lengths data used in this paper, as well as the scripts used to perform the analysis are available at our Zenodo repository via <https://doi.org/10.5281/ZENODO.13974204> (Zou & Fialko, [2024\)](#page-15-0).

The original Nevada fault map data used for concatenation and lengths calculation is available at the US Quaternary Fault and Fold Database via [https://www.usgs.gov/programs/earthquake‐hazards/faults](https://www.usgs.gov/programs/earthquake-hazards/faults) (U.S. Geological Survey and Nevada Bureau of Mines and Geology, [2023\)](#page-14-0).

The original Ventura County fault map data used for concatenation and lengths calculation is available at the Ventura County Faults database via [https://venturacountydatadownloads‐vcitsgis.hub.arcgis.com/datasets/](https://venturacountydatadownloads-vcitsgis.hub.arcgis.com/datasets/2e356ea3e3df4c0bbfb8cfd77681bf20_0/explore?location=34.423653%2C-119.059450%2C8.90) [2e356ea3e3df4c0bbfb8cfd77681bf20_0/explore?location](https://venturacountydatadownloads-vcitsgis.hub.arcgis.com/datasets/2e356ea3e3df4c0bbfb8cfd77681bf20_0/explore?location=34.423653%2C-119.059450%2C8.90)=34.423653%2C‐119.059450%2C8.90 (County of Ventura, [2023\)](#page-12-0).

The original Pennsylvania fault map data used for concatenation and lengths calculation is available at the PA Bedrock Geology Faults (Vector) database via https://newdata-dcnr.opendata.arcgis.com/datasets/DCNR:: bedrock-geology-of-pennsylvania-faults-vector/about (PA Department of Conservation & Natural Resources, [2023](#page-13-0)).

The original New Jersey fault map data used for concatenation and lengths calculation is available at the Geologic Faults in New Jersey database via [https://www.arcgis.com/home/item.html?id](https://www.arcgis.com/home/item.html?id=3a5c8b0a00be4485bcd22281a76f8e23)=3a5c8b0a00be4485bcd2 [2281a76f8e23](https://www.arcgis.com/home/item.html?id=3a5c8b0a00be4485bcd22281a76f8e23) (NJ Dept. of Environmental Protection Bureau of GIS, [2023\)](#page-13-0).

The original New Zealand fault map data used for concatenation and lengths calculation is available at the New Zealand Active Faults Database: High Resolution via <https://data.gns.cri.nz/af/> (Langridge et al., [2016b](#page-13-0)).

The Southern California earthquake catalog used for the analysis of the cumulative size-frequency distribution of earthquakes in Southern California recorded over the last 42 years is available at the 1981–2023 Catalog via [scedc.caltech.edu/data/alt‐2011‐dd‐hauksson‐yang‐shearer.html](https://scedc.caltech.edu/data/alt-2011-dd-hauksson-yang-shearer.html) (Hauksson et al., [2024](#page-13-0)).

The global CMT catalog for the analysis of the magnitude‐frequency distribution of strike‐slip earthquakes that occurred between 1976 and 2024 is available at the Global CMT Catalog Search via www.globalcmt.org (Ekström et al., [2024\)](#page-12-0).

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